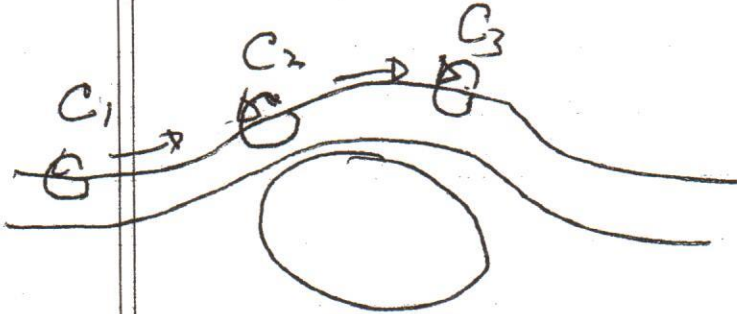


## (c) Potential Flow

- Consider fluid streamlines:



if  $\underline{\omega} = 0$  at any point along streamline, then Kelvin's thm  $\Rightarrow \underline{\omega} = 0$  everywhere on streamline.

Easily seen by considering circulation around infinitesimal loop "pulled" along streamline. Thus, if

$$\oint_{C_1} \underline{v} \cdot d\underline{l} = \int_{A_1} \underline{\omega} \cdot d\underline{s} = 0, \text{ then } \oint_{C_n} \underline{v} \cdot d\underline{l} = \int_{A_n} \underline{\omega} \cdot d\underline{s} = 0$$

for all  $C_n$ .

- flow with  $\underline{\omega} = \nabla \times \underline{v} = 0$  in all space is defined as:

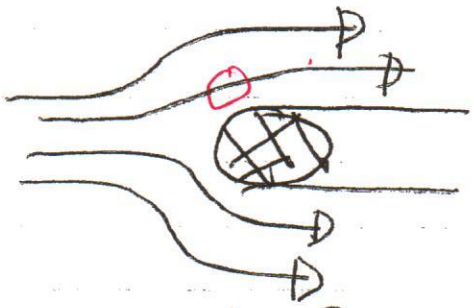
$\Rightarrow$  potential, irrotational flow

$\Leftarrow \underline{\omega} \neq 0$  rotational, vortical flow

- Important to note breakdown of Kelvin's Thm

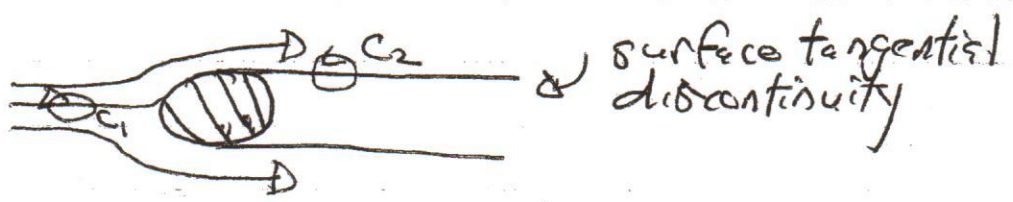
Applicability, namely to flows with separation

c.e. consider flow around sphere



- c.e. - streamlines separate from body
- surface of tangential discontinuity appears (velocity component)
- ⇒ - Kelvin Thm not applicable

c.e.



- cannot infer  $\oint_{c_2} \underline{v} \cdot d\underline{l}$  from  $\oint_{c_1} \underline{v} \cdot d\underline{l}$  due to separation-induced tangential discontinuity
- Also, viscosity important in (boundary layer) region of discontinuity. As viscous effects  $\sim \nu k^2$ , deviation from potential flow naturally most significant in small scale region of boundary layer!

Now, for isentropic fluids:

$$\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \nabla \underline{v} = -\nabla W$$

$W \equiv$  enthalpy

stream function

for potential flow,  $\underline{v} = \underline{\nabla}\phi \Rightarrow \underline{\omega} = 0$

$$\underline{v} \cdot \underline{\nabla} \underline{v} = -\underline{v} \times \underline{\nabla} \phi + \underline{\nabla} (v^2/2)$$

$$= 0 + \underline{\nabla} (v^2/2), \text{ for potential flow}$$

$$\frac{\partial \underline{v}}{\partial t} + \underline{\nabla} (v^2/2) = -\underline{\nabla} W$$

$$\underline{v} = \underline{\nabla} \phi$$

$$\Rightarrow \underline{\nabla} \left( \frac{\partial \phi}{\partial t} + \frac{(\underline{\nabla} \phi)^2}{2} + W \right) = 0$$

∴ have equation for dynamics of potential flow:  
*Bernoulli* *along streamlines*

$$\frac{\partial \phi}{\partial t} + \frac{(\underline{\nabla} \phi)^2}{2} + W = f(t)$$

$f(t)$  defined for each streamline

- for  $\partial \phi / \partial t = 0$ , recover Bernoulli's Law

- obvious that potential not uniquely defined, as  $\underline{v} = \underline{\nabla} \phi$



what does incompressibility  
mean?

26

Now, consider incompressible fluid potential flow,

- flows leaving density constant (no compression, expansion)

$$\underline{\nabla \cdot \underline{v}} = 0 \quad \Leftrightarrow \quad \frac{d\rho}{dt} = 0 \quad (\rho \text{ constant})$$

$$\Rightarrow \text{if } \underline{v} = \underline{\nabla} \phi \quad \Rightarrow \quad \boxed{\nabla^2 \phi = 0}$$
$$\frac{\partial \phi}{\partial t} + \frac{(\underline{\nabla} \phi)^2}{2} + \frac{p}{\rho} = f(t)$$

$\therefore$  for static flow, with gravity,  $\Rightarrow$  Bernoulli Eqn.:

$$\boxed{v^2/2 + p/\rho + gz = \text{const.}}$$

Criterion for "incompressibility":

- "incompressibility" is valid description for certain classes of flows, dependent on time scales, speeds, etc.

- for stationary flows

$$\underline{\partial \underline{v} / \partial t} = 0,$$

$$\frac{p}{\rho} + \frac{v^2}{2} = \text{const.}$$

Now, for adiabatic fluid (T const)  $\Rightarrow$   
isentropic

$$\Delta P = \left( \frac{\partial P}{\partial \rho} \right)_s \Delta \rho$$

but  $\rho + \frac{v^2}{2} = \text{const.}$

$$\Rightarrow \Delta \left( \frac{v^2}{2} \right) = - \left( \frac{\partial P}{\partial \rho} \right)_s \frac{\Delta \rho}{\rho}$$

"Incompressibility"  $\Rightarrow \Delta P / \rho \ll 1$

$$\left( \frac{\partial P}{\partial \rho} \right)_s = c_s^2 \quad (\text{sound speed in fluid})$$

$\therefore v^2 / c_s^2 \ll 1 \Rightarrow$  Flow incompressible

Note:  $\left\{ \begin{array}{l} \text{Supersonic flows always compressible} \Rightarrow \\ \text{fluid dynamics coupled to acoustic waves} \end{array} \right.$

- for dynamic flows (more generally);

need compare terms in continuity equation

$$\frac{d\rho}{dt} = -\rho \nabla \cdot \underline{v}$$

Now  $\tau \rightarrow$  time scale for flow  
 $l \rightarrow$  spatial scale for flow

Then,  $\frac{dp}{dt} \sim \frac{\Delta p}{\tau}$

$\rho \nabla \cdot \underline{v} \sim \rho \frac{\tilde{v}}{l}$

$\nabla \cdot \underline{v} \sim 0$   
 $\frac{\partial \underline{v}}{\partial t} \sim \frac{\Delta p}{\tau}$   
 $\frac{\partial \underline{v}}{\partial t} \sim \frac{\Delta p}{\tau} \cdot \frac{1}{c_s^2}$

To relate  $\Delta p$  to  $\tilde{v}$ , consider Euler equation:

$\frac{\partial \underline{v}}{\partial t} = - \frac{\nabla p}{\rho}$

$\rho \frac{\tilde{v}}{l} > \pm \frac{\rho l}{\tau} \frac{\rho l}{c_s}$   
 $c_s^2 > l^2 / \rho \tau$

$\Rightarrow \tilde{v} \sim \frac{c_s^2 \Delta p}{\rho l}$

(typically,  $\tilde{v} \approx l/\tau$ )

$\tilde{v} \sim \tau c_s^2 \Delta p$

$\frac{dp}{dt} \sim \left( \frac{c_s^2 \rho}{l} \right) \tilde{v}$

$\rho \nabla \cdot \underline{v} \sim \frac{\tilde{v}}{l} \rho$

$\rho \nabla \cdot \underline{v} \sim \frac{\rho l}{\tau}$   
 $\rho \frac{\tilde{v}}{l} \sim \frac{\Delta p}{\tau}$   
 $\frac{\rho l}{\tau} \sim \frac{\rho l}{\tau} \frac{c_s^2 \Delta p}{\rho l}$   
 $\sim \frac{c_s^2 \Delta p}{\tau}$

$\nabla \cdot \underline{v} \approx 0$  if  $\frac{dp}{dt} \ll \rho \nabla \cdot \underline{v}$



$$\frac{\tilde{v}}{l} \rho \gg \frac{\rho}{c_s^2 \gamma^2} \tilde{v} \Rightarrow c_s^2 \gg \frac{l^2}{\gamma^2}$$


Thus, dynamics is compressible if  $\begin{cases} c_s^2 \gg l^2/\gamma^2 \\ \gg \omega^2/k^2 \text{ (wave)} \end{cases}$

- note can synthesise static, dynamic conditions to obtain incompressibility criterion:

$$c_s^2 > \begin{cases} \tilde{v}^2 \\ l^2/\gamma^2 \end{cases} \quad \text{i.e. } \begin{cases} \text{time slow compared to} \\ \text{time to traverse 1 spatial} \\ \text{scale at acoustic speeds.} \end{cases}$$

Some further facts about potential flows (generally incompressible):

- for body (i.e. rigid sphere) immersed in fluid, if amplitude oscillation  $\ll$  dimensions of body  $\Rightarrow$  motion describable by potential flow

i.e.  $q \equiv$  amplitude motion   
 $u \equiv$  body velocity  
 $f \equiv$  frequency of oscillation  
 $l \equiv$  size of body

Simply compare  $\frac{\partial v}{\partial t}$  to  $\underline{v \cdot \nabla v}$ , noting

OR  $\frac{\partial v}{\partial t} \gg \underline{v} \cdot \underline{\nabla} v$

30

$$\text{if } \frac{\partial v}{\partial t} \gg \underline{v} \cdot \underline{\nabla} v \Rightarrow \frac{\partial v}{\partial t} \equiv -\underline{\nabla} w$$

$$\text{so } \underline{\nabla} \times \underline{v} = 0 \Rightarrow \left\{ \begin{array}{l} \text{Potential} \\ \text{flow} \end{array} \right.$$

Now  $w \sim u/a$

$$\frac{\partial v}{\partial t} = -i\omega \underline{v} \sim u^2/a \quad (\underline{v} \sim u \text{ near body})$$

$$\underline{v} \cdot \underline{\nabla} v \sim u^2/l \quad (l \text{ sets smallest scale in problem})$$

$$\left| \frac{\partial v}{\partial t} \right| \gg \left| \underline{v} \cdot \underline{\nabla} v \right| \Rightarrow \frac{u^2}{a} \gg \frac{u^2}{l}$$

$$\Rightarrow l \gg a$$

Thus, fluid dynamics resulting from small oscillation of body describable by potential flow.

- In potential flow, streamlines must be open, not closed.

To see, consider circulation about closed contour



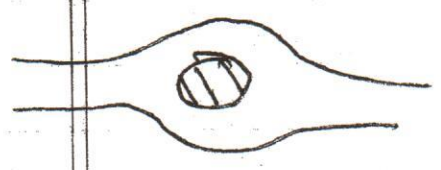
$\phi$  changes

$$\oint_C \underline{v} \cdot d\underline{l} = \int_{S_{enc.}} d\underline{s} \cdot \underline{\omega} = 0$$

$\underline{\omega} = 0$  for potential flow

but, by definition,  $\int_{streamline} \underline{v} \cdot d\underline{l} \neq 0 \Rightarrow$  streamlines must be open!

c.e.



sphere in  $\underline{v} = v_0 \hat{z}$  flow is typical potential flow problem (describes flow at distance from sphere).

- For incompressible flow, (not potential)

$$\frac{d\underline{\omega}}{dt} = \underline{\omega} \cdot \nabla \underline{v} - \underline{\omega} \nabla \cdot \underline{v}$$

In 2D,  $\underline{\omega} \cdot \nabla \underline{v} = 0$  i.e.  $\begin{cases} \underline{v} = (v_x(x,y), v_y(x,y)) \\ \underline{\omega} = \omega_z(x,y) \hat{z} \end{cases}$

Then,  $\frac{d\underline{\omega}}{dt} = 0$

Now,  $\nabla \cdot \underline{v} = 0 \Rightarrow \begin{cases} v_x = \partial\psi/\partial y \\ v_y = -\partial\psi/\partial x \end{cases}$

$$\underline{\omega} = \left( \frac{\partial v_y}{\partial x} - \frac{\partial v_x}{\partial y} \right) \hat{z} = \hat{z} (-\nabla^2 \psi)$$

$$\frac{d\underline{\omega}}{dt} = 0 \Rightarrow \begin{cases} + \frac{\partial}{\partial t} \nabla^2 \psi + \underline{\nabla} \psi \times \underline{z} \cdot \nabla \nabla^2 \psi = 0 \\ \text{2D incompressible fluid eqn.} \end{cases}$$

iv.) Problems in Potential Flow

a.) Incompressible Potential Flow Around Sphere

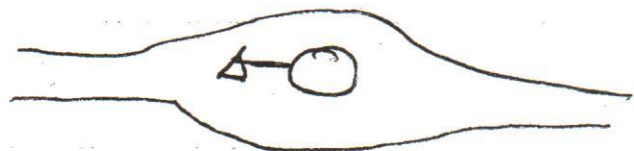
Consider <sup>rigid</sup> sphere in motion at  $\underline{u}$  in infinite fluid



Flow Pattern ?

Now :

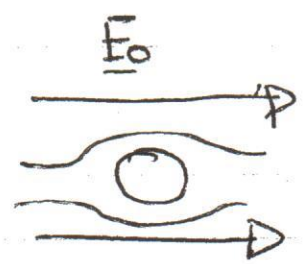
- intuitively, expect :



i.e. equivalent to  $\begin{cases} \text{sphere at rest} \\ \underline{v}|_{\text{fluid}} = -\underline{u} \\ \infty \end{cases}$

Electrostatic analogy: Conducting sphere in uniform electric field

i.e.



$$\phi = -\underline{E}_0 \cdot \underline{r} + \phi_{\text{sphere}}$$

$\phi_{\text{sphere}}$  is dipole field.

Dipole moment determined by b.c.

i.e.  $\phi = \text{const} = 0$  on sphere surface

Now, for potential flow (incompressible):

$$\nabla^2 \phi = 0$$

$$\underline{v} = \nabla \phi$$

$$v_n = \underline{v} \cdot \hat{n} = u \cdot \hat{n} \Big|_{\text{surface}}$$

(i.e. normal velocity = sphere velocity on surface)


By analogy with electrostatics, can solve via:

- multipole expansion
- b.c.'s determine effective "charge" distribution



Recall e.s.  $\Rightarrow \nabla^2 \phi = -4\pi\rho$

$$\phi = \int d^3x' \frac{\rho(x')}{|x-x'|}$$

For  $\underline{x}$  outside region  $\rho$  : 

$$\phi(\underline{x}) = \int d^3x' \frac{\rho(x')}{|x-x'|}$$

$$= \int d^3x' \frac{\rho(x')}{|x-x'|} = \int d^3x' \underline{x}' \rho(x') \cdot \nabla \left( \frac{1}{|\underline{x}|} \right) + \dots$$

$$= \underbrace{\frac{Q}{|\underline{x}|}}_{\text{monopole}} - \underbrace{\underline{d} \cdot \nabla \left( \frac{1}{|\underline{x}|} \right)}_{\text{dipole}} + \dots \underbrace{\quad}_{\text{quadrupole}}$$

Thus, can write down general solution for potential flow streamlines around body as multipole expansion.


$\rightarrow Q = 0$  (no sources, sinks)

$\therefore$  in general dipole dominates

→ in 2D, same story with  $\ln|x-x'| \rightarrow 1/|x-x'|$

Here:  $\underline{u} = u \hat{z}$  (slow velocity) (spherical symmetry) (body velocity)

$V_n|_R = V_r|_R = u \hat{z} \cdot \hat{n} = u \cos\theta$  } boundary condition



Now,  $\phi(\underline{x}) = \underline{A} \cdot \underline{\nabla} (1/|\underline{x}|)$  u → z.

$\underline{A} = A \hat{z}$  (dipole moment in  $\hat{z}$  direction)

$\phi = -A \frac{\cos\theta}{r^2}$

$V_r = 2A \cos\theta / r^3$

$V_r = u \cos\theta$   
on sphere

$\Rightarrow \frac{2A \cos\theta}{R^3} = u \cos\theta$

$\Rightarrow A = \frac{R^3}{2} u$

$\phi = -u R^3 \cos\theta / 2 r^2$

$\underline{v} = \underline{\nabla} \phi$

determined general flow field

Note:

- can recover from  $\phi = \sum \left( \frac{a_n}{r^n} + \frac{b_n}{r^{n+1}} \right) P_n(\cos \theta)$  <sup>regularity at  $\infty$</sup>   
expansion and b.c.'s.

- if sphere in uniform field:

$$\phi = U_0 r \cos \theta + \phi_{\text{sphere}}$$

$\phi$  determine from  $V_n = 0$

to determine pressure distribution on sphere,

Recall:  $\rho \frac{\partial \phi}{\partial t} + \frac{\rho v^2}{2} + p = p_0$  } incompressible Bernoulli Eqn.  
 $\downarrow$   
 ambient pressure at  $\infty$

Thus, can immediately write:

$$p(x) = p_0 - \frac{\rho}{2} \nabla \phi \cdot \nabla \phi - \rho \frac{\partial \phi}{\partial t}$$

$\phi(x) \equiv$  determined at  $\infty$  above via  $\nabla^2 \phi = 0$   
and b.c.'s.



As sphere in motion (but uniform) :

$$\frac{\partial \phi}{\partial t} = -\underline{u} \cdot \nabla \phi + \frac{\partial \phi}{\partial y} \dot{y}$$

$\dot{y} = 0$

so

$$P(x) = P_0 - \frac{\rho}{2} \nabla \phi \cdot \nabla \phi - \underline{u} \cdot \nabla \phi$$

Generally, leads to concept of stagnation point

i.e. for Bernoulli Eqn. for incompressible fluid :

$$\frac{P}{\rho} + \frac{V^2}{2} = \text{const.} = P_0$$

Now, consider fixed body in fluid with  $\begin{cases} V_{\infty} = U_0 \\ P_{\infty} = P_0 \end{cases}$

As  $V = 0$  on surface body :

$$P_{\text{max}} = P|_{\text{bdy}} = P_0 + \frac{1}{2} \rho U^2$$

- stagnation point ( $V=0$ ) on body is point of maximal pressure

- maximal pressure determined by  $\begin{cases} P_0 \\ \text{speed} \end{cases}$



→ Fish skeleton strongest on front face, weakest elsewhere

→ front face is point of maximal pressure ('head')

→ eye lens adjusts to allow for speed-induced pressure changes.

### b.) Drag Force and Induced Mass

bubble

→ Heuristics: Consider rigid body in water.

→ what force cell does parallel u?



Slow body motion ⇒ potential flow around sphere  
⇒ energy in fluid motion, too!

Thus, for F<sub>ext</sub> to move body in fluid, need work against  
- inertia of body (obvious)  
- inertia of fluid, excited into potential flow

Thus, for body in water, need interpret Newton's 2nd Law as:

$$\underline{F}_{ext} = M_{eff} \frac{d\underline{y}}{dt}$$

$M_{eff} = M + m_{induced}$  → induced mass of fluid in potential flow around body  
↓ (mass of fluid flow which addresses the body)  
mass of body in water simplest form possible

To calculate induced mass:

- ⊕ - calculate energy in potential flow around rigid body in uniform motion in fluid
- ⊕ - use  $dE = d\underline{p} \cdot \underline{y}$  to determine momentum in fluid

as  $\underline{p} = \underline{p}(\underline{y}) \Rightarrow \underline{p}_i = m_{ik} U_k$

∴  $m_{ik}$  is induced mass tensor!

→ Calculation: Consider rigid body moving in fluid

i.e.



Now, for flow field outside body, multipole expansion solution to  $\nabla^2 \phi = 0$  yields



$$\phi = \frac{\phi}{r} + \underline{A} \cdot \underline{\nabla} \left( \frac{1}{r} \right) + \dots$$

$\uparrow$   
 monopole  
 (vanishes  $\rightarrow$   
 no sources)

$\uparrow$   
 dipole  
 (dominant multipole  
 at large radius)

$\rightarrow$  dipole moment:  $A = c R^3 \underline{u}$

$\therefore \phi = \underline{A} \cdot \underline{\nabla} (1/r)$  ( $c = 1/2$ , sphere)

$$= - \underline{A} \cdot \underline{r} / r^3 = - \underline{A} \cdot \underline{\hat{n}} / r^2$$

$$\underline{v} = \underline{\nabla} \phi = \underline{A} \cdot \underline{\nabla} \nabla (1/r)$$

$$= (\underline{A} \cdot \underline{\nabla}) (-1/r^3)$$

$$\underline{v} = (3(\underline{A} \cdot \underline{\hat{n}}) \underline{\hat{n}} - \underline{A}) / r^3$$

$$\phi = -A \frac{\cos \theta}{r^2}$$

$$v_r = \frac{2A \cos \theta}{r^3}$$

$$v_{\theta} =$$

$$\frac{2A \sin \theta}{r^3}$$

$$A = \frac{4}{3} R^3$$

Now, for energy, seek calculate fluid energy in volume  $V$  enclosed within radius  $R$  around body. Take  $R^3 \gg V_0 \equiv$  volume of body.

Thus:

$$E = \frac{1}{2} \rho \int dV |\underline{v}|^2$$

$$= \frac{1}{2} \rho \int d^3x (\underline{u}^2 + |\underline{\hat{n}}|^2 - \underline{u}^2)$$

$$\begin{aligned}
 \text{out } \nabla^2 u^2 &= (\underline{v} + \underline{u}) \cdot (\underline{v} - \underline{u}) \\
 &= \nabla(\phi + \underline{u} \cdot \underline{r}) \cdot (\underline{v} - \underline{u}) \\
 &= \nabla \cdot [(\phi + \underline{u} \cdot \underline{r})(\underline{v} - \underline{u})]
 \end{aligned}$$

$$\begin{aligned}
 \text{as } \underline{v} &= \nabla \phi & \nabla \cdot \underline{v} &= 0 \\
 \underline{u} &= \text{const.} & \nabla \cdot \underline{u} &= 0
 \end{aligned}$$

$$\therefore E = \frac{1}{2} \rho \int d^3x \left[ u^2 + \nabla \cdot [(\phi + \underline{u} \cdot \underline{r})(\underline{v} - \underline{u})] \right]$$

$$= \frac{1}{2} \rho u^2 (V - V_0) + \frac{1}{2} \rho \int d\underline{s} \cdot [(\phi + \underline{u} \cdot \underline{r})(\underline{v} - \underline{u})]$$

$\int$  volume space  $\rightarrow$  Volume object/body

$$V = \frac{4\pi}{3} R^3$$



$$\begin{cases} (\underline{v} - \underline{u}) \cdot d\underline{s} = 0 \\ \text{on } R_0 \text{ surface} \end{cases}$$

$$\text{Now, } d\underline{s} = \underline{\hat{n}} R^2 d\Omega, \text{ on outer surface}$$

$$E = \frac{1}{2} \rho u^2 (V - V_0)$$

$$+ \frac{1}{2} \rho \int R^2 d\Omega [(\underline{\hat{n}} \cdot \underline{v} - \underline{\hat{n}} \cdot \underline{u})(\phi + \underline{u} \cdot \underline{r})]$$

$$E = \frac{1}{2} \rho u^2 (V - V_0) + \frac{1}{2} \rho \int R^2 d\Omega \left[ \left( 2 \frac{(\underline{A} \cdot \underline{n})}{R^3} - \underline{u} \cdot \underline{n} \right) \left( -\frac{\underline{A} \cdot \underline{n}}{R^2} + R \underline{u} \cdot \underline{n} \right) \right]$$

$$= \frac{1}{2} \rho u^2 (V - V_0) + \frac{1}{2} \rho \int R^2 d\Omega \left[ -2 \frac{(\underline{A} \cdot \underline{n})^2}{R^5} \right] \quad \text{vanishes for large } R$$

$$+ \left[ \frac{(\underline{u} \cdot \underline{n})(\underline{A} \cdot \underline{n})}{R^2} + \frac{2(\underline{A} \cdot \underline{n})(\underline{u} \cdot \underline{n})}{R^2} - R (\underline{u} \cdot \underline{n})^2 \right]$$

$$= \frac{1}{2} \rho u^2 (V - V_0) + \frac{1}{2} \rho \int R^2 d\Omega \left[ \frac{3(\underline{A} \cdot \underline{n})(\underline{u} \cdot \underline{n})}{R^2} - R^3 (\underline{u} \cdot \underline{n})^2 \right]$$

Thus finally,

$$E = \frac{1}{2} \rho u^2 (V - V_0) + \frac{1}{2} \rho \int d\Omega \left[ 3(\underline{A} \cdot \underline{n})(\underline{u} \cdot \underline{n}) - R^3 (\underline{u} \cdot \underline{n})^2 \right]$$

$$d\Omega = d\theta \sin\theta d\phi$$

$$\text{if } \int d\Omega ( ) = \langle ( ) \rangle$$

$$\Rightarrow \langle (\underline{A} \cdot \underline{n})(\underline{B} \cdot \underline{n}) \rangle = \frac{1}{2} \delta_{ij} A_i B_j = \frac{1}{3} \underline{A} \cdot \underline{B}$$



$$E = \frac{1}{2} \rho u^2 (V - V_0) + \frac{1}{2} \rho \left[ 4\pi \underline{A} \cdot \underline{y} - \frac{4\pi}{3} R^3 u^2 \right]$$

$$= \frac{1}{2} \rho \left[ 4\pi \underline{A} \cdot \underline{y} - u^2 V_0 \right]$$

Thus finally,

$$E = \frac{\rho}{2} \left[ 4\pi \underline{A} \cdot \underline{y} - u^2 V_0 \right]$$

energy in potential flow around body

Now,  $\underline{A} = \underline{A}(u) \Rightarrow \left\{ \begin{array}{l} E = \frac{1}{2} m_{ik} u_i u_k \\ \text{defines induced mass tensor} \end{array} \right.$

$$dE = \underline{y} \cdot d\underline{P}$$

$$\Rightarrow \underline{P} = \frac{\rho}{4\pi} \left[ 4\pi \underline{A} - V_0 \underline{y} \right]$$

momentum in potential flow

Now, consider external force acting system, where system = body + fluid (in Pot. flow)

i.e.  $\underline{F}_{ext} = \frac{dP_{fluid}}{dt} + M_{body} \frac{d\underline{U}}{dt}$

$\Rightarrow \underline{f}_k = (M \delta_{ik} + m_{ik}) \frac{d\underline{U}_k}{dt}$

∴ effective mass of "system" is sum of - body mass

- induced mass of fluid in potential flow around body

→ note induced mass is determined purely by body shape (i.e. via volume and dipole moment)

i.e. for sphere  $\underline{A} = \frac{R_0^3}{2} \underline{U}$

$$\underline{P} = \rho \left[ 4\pi \frac{R_0^3}{2} \underline{U} - \frac{4\pi}{3} R_0^3 \underline{U} \right]$$

$$= \rho \frac{2}{3} \pi R_0^3 \underline{U}$$

$$m_{induced} = \rho \frac{2}{3} \pi R_0^3$$

In general  $M_{induced} \sim \# \rho R^3$

$\sim \# \rho V$   
 $\downarrow$   $\rightarrow$  displaced mass  
 Numerical fluid  
 factor, shape dependent

$\rightarrow$  Example of "renormalization" in classical physics "dressing field" in continuum i.e. {renorm. polarization, Debye shield, etc}

i.e. in quantum electrodynamics  $\rightarrow$  electron polarizes vacuum

$\rightarrow m_e = m_e^{bare} + m_e^{V.P.}$   
 $(E=mc^2)$

in classical potential flow  $\rightarrow$  moving a sphere in  $H_2O$  requires that some energy go into surrounding media (the water!)

(skip)

$\rightarrow$  Enhanced inertia due induced mass may alternatively, be viewed as drag force on body (careful of phases!)  
 mom. transmittal to fluid

i.e.  $F_{ext} = \frac{dP_{fluid}}{dt} + M \frac{dy}{dt}$



$$M \frac{dy}{dt} = \underbrace{f_{ext}} - \underbrace{\frac{dP_{fluid}}{dt}}_{\substack{\text{drag!} \\ \downarrow}} \\ = \underbrace{f_{ext}} + \underbrace{f_{drag, lift}} \quad f_{drag} \sim u$$

$f_{drag} = -\frac{dP_{fluid}}{dt}$ , along direction motion.

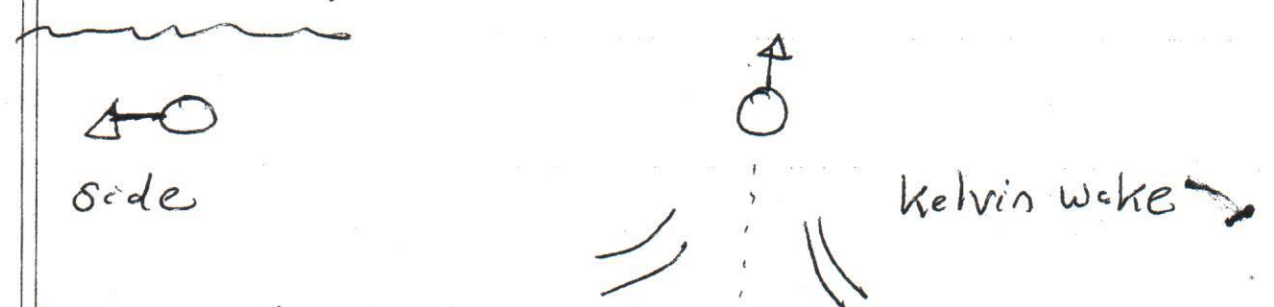
$f_{lift} = -\frac{dP_{fluid}}{dt}$ ,  $\perp$  direction of motion.

Note:  $\rightarrow$  if body is uniform motion in ideal (fantasy) fluid  $f_{drag} = f_{lift} = 0$  } D'Alembert's paradox

$\rightarrow$  need external force to maintain uniform motion

as no - no dissipation (ideal fluid)  
 - no loss of energy to  $\infty$  ( $V \sim 1/R^3$ )

$\rightarrow$  but if body near surface



body will radiate surface waves to  $\infty$  (wake)  $\Rightarrow$  wave drag induced energy loss!



→ Related Problem:

- consider body in fluid, which is set in motion by external agent



Relate  $\underline{u}$  body to  $\underline{v}$  fluid!?

- Now  $\underline{v} \equiv$  velocity of unperturbed flow

$$\frac{\|\nabla \underline{v}\|}{\|\underline{v}\|} R_0 \ll 1 \Rightarrow \underline{v} \sim \text{const over scale of body} \\ (\text{potential flow valid})$$

so if body fully carried along by fluid ( $\underline{v} = \underline{u}$ ), then force on it would equal force on volume of displaced fluid

i.e. 
$$\frac{d}{dt} (M \underline{u}) = \rho V_0 \frac{d\underline{v}}{dt}$$

but body moves relative to fluid, so that fluid acquires momentum → due to relative motion

i.e. 
$$\frac{d\underline{p}_{\text{fluid}}}{dt} = -\underline{m} \cdot \frac{d}{dt} [\underline{u} - \underline{v}]$$



∴ so really,

$$\frac{d}{dt} (M \underline{u}) = \rho V_0 \frac{d\underline{v}}{dt} = \frac{d}{dt} (m (\underline{u} - \underline{v}))$$

$$\frac{d}{dt} (M u_i) = \rho V_0 \frac{d v_i}{dt} = m_{ik} \frac{d}{dt} (u_k - v_k)$$

⇒

$$M u_i = \rho V_0 v_i - m_{ik} (u_k - v_k)$$

$$(M \delta_{ik} + m_{ik}) u_k = (\rho V_0 \delta_{ik} + m_{ik}) v_k$$

$$u_k = \frac{(\rho V_0 \delta_{ik} + m_{ik})}{M \delta_{ik} + m_{ik}} v_k$$

Note:  $\rho V_0 < M$  (body heavier than displaced fluid) → body lags

$\rho V_0 > M$  → body leads

$\rho V_0 = M$   $u_k = v_k$ .

Thus

$$M \frac{du}{dt} = \rho_f V \frac{dv}{dt} - \underline{m} \cdot \frac{d}{dt} [u - v]$$

$$(M \delta_{ij} + m_{ij}) \frac{du_j}{dt} = M_f \delta_{ij} + m_{ij} \frac{dv_j}{dt}$$

$$\therefore u_j = \left[ (M_f \delta_{ij} + m_{ij}) / (M \delta_{ij} + m_{ij}) \right] v_j$$

$$M_f = \rho_f V_0$$

$$M = \rho V_0$$

$$\Rightarrow \underline{u} = \underline{v} \quad \text{if} \quad \rho_f = \rho$$

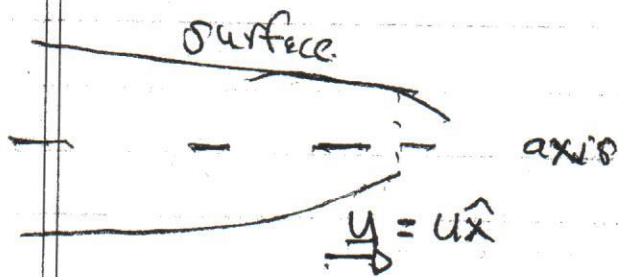
$$u < v \quad \text{if} \quad \rho_f < \rho \quad \rightarrow \text{heavy object} \\ \text{lags}$$

$\rho_f \equiv$  fluid density  
 $\rho \equiv$  body density

$$u > v \quad \text{if} \quad \rho_f > \rho \quad \rightarrow \text{light object lags}$$

### c.) Potential Flow - General Slender Body

- Till now, have considered simple body potential flows, i.e. sphere, cylinder,  
 Here consider general body from surface of revolution



- i.e.
- generally axially symmetric slender body
  - slender  $\Rightarrow w/L \ll 1$

Now, observe analogy with electrostatics again,

i.e. e.s.  $\Rightarrow \phi(x) = \int d^3x' \rho(x') / |\underline{x} - \underline{x}'|$

potential flow ( $A \sim uV$ )

$$\phi(x) = \frac{1}{4\pi} \int d^3x' (\dot{\rho}(x') / \rho_0) / |\underline{x} - \underline{x}'|$$

$\frac{\dot{\rho}(x')}{\rho_0}$   $\equiv$  normalized density of fluid flowing across cross-section of body

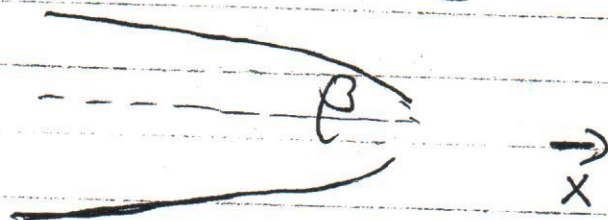
$\rightarrow$  yields  $A \sim V_0 U$  etc.

$$\phi(x) = \frac{1}{4\pi |\underline{x}|^2} \int d^3x' \frac{\dot{\rho}(x')}{\rho_0} \underline{x}' + \text{h.o.t.}$$

$\downarrow$   
dipole term dominates



Flow, body slender  $\rightarrow \frac{W}{L} \ll 1 \Rightarrow \beta \ll 1$



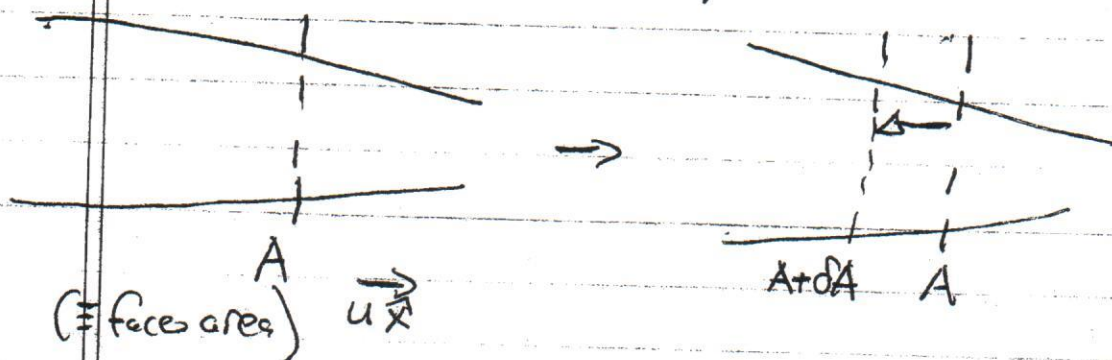
$\nabla \cdot \underline{V} = 0$  and axial symmetry  $\Rightarrow$

$$\frac{\partial V_x}{\partial x} + \frac{1}{r} \frac{\partial}{\partial r} (r V_r) = 0$$

$$\therefore \frac{V_r}{V_x} \sim \frac{\Delta r}{\Delta x} \sim \beta \sim \frac{W}{L} \ll 1$$

$\Rightarrow$  need only consider  $\vec{x}$  fluid motion

$\therefore$  to compute dipole moment, need  $\rho(x)/\rho_0$  for fluid flow across body



$$\text{Net } \frac{\dot{p}}{\rho_0} = u \left[ A + \delta A - A \right] = u \frac{\partial A}{\partial x} dx$$

$$\Rightarrow \rho(x')/\rho_0 = u \frac{\partial A}{\partial x'}$$

$$\therefore \phi(x) = \frac{1}{4\pi/x^2} \int dx' x' u \frac{\partial A(x')}{\partial x'}$$

$$= \frac{-u}{4\pi/x^2} \int dx' A(x')$$

$$= \frac{-u V}{4\pi/x^2}$$

$$V \equiv \text{volume of body} = \int dx' A(x')$$

$\Rightarrow$  yields intuitive result:

$$\phi(x) = \frac{-u V_{\text{body}}}{4\pi r^2}$$

effective dipole moment for slender body.